



Search for black holes using gravitational microlensing events observed by Gaia

Katarzyna Kruszyńska, Łukasz Wyrzykowski, P. Zieliński, K. A. Rybicki, M. Gromadzki, M. Jabłońska, K. Howil, N. Ihane et al. (Astronomical Observatory, University of Warsaw)

Seminarium Ogólne, UMK, Toruń, 30.05.2022

Outline

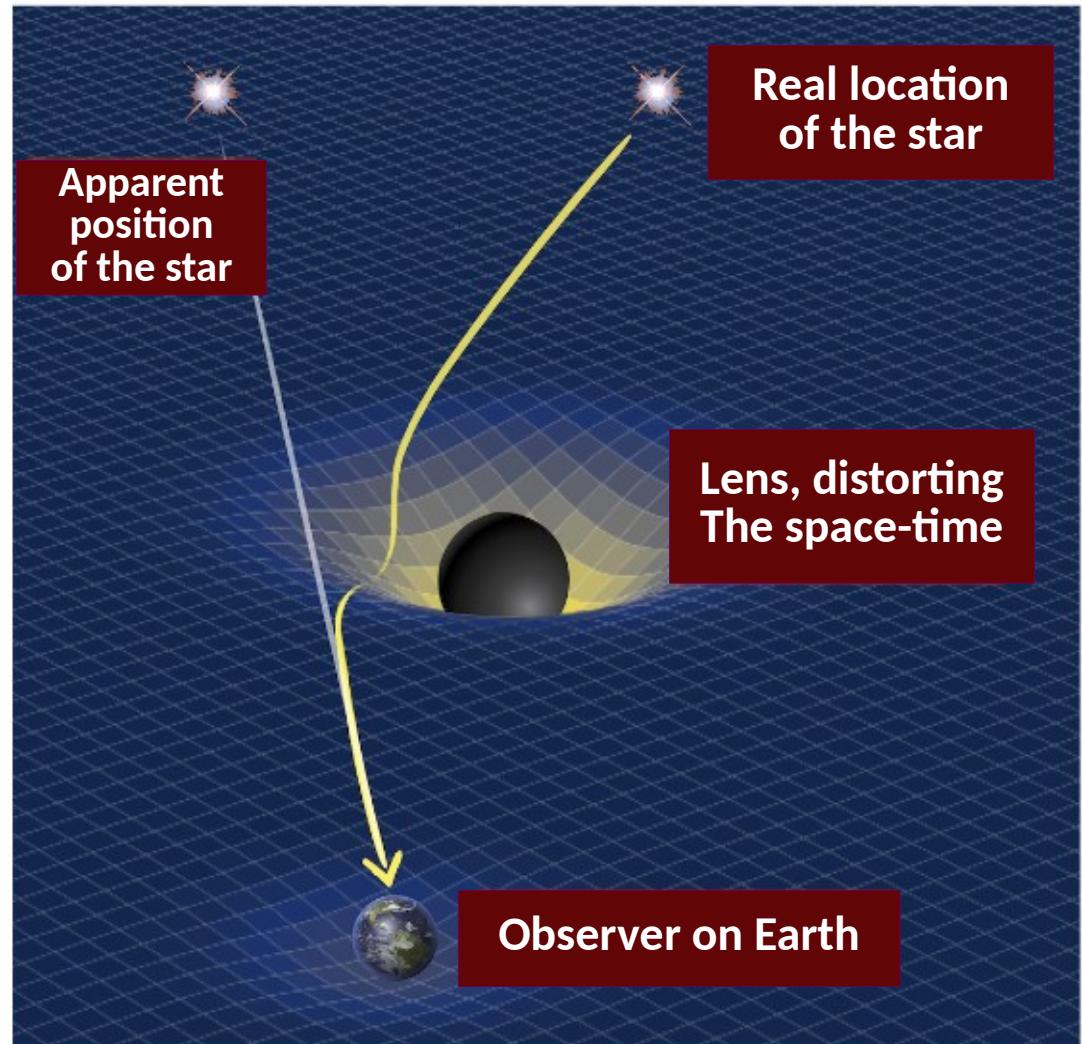
- 1) Black holes – known and unknown**
- 2) Gravitational microlensing – history, use, surveys and future**
- 3) Gaia and Gaia Science Alerts**
- 4) Gaia and Microlensing**
- 5) Constraining mass and distance to the lens explained on Gaia18cbf**
- 6) The Future**

Black holes (and other dark remnant): What Do We Know? Do We Know Things?? Let's Find Out!

- Distribution of remnants: where are they? Are BHs in Galactic bulge or disk or halo?
- How massive are the remnants and what is their mass distribution?
- Mass gap: does it exist or not?
- Natal kick velocity of remnants? Is it large or small?
- Source of knowledge:
 - LIGO-VIRGO merger rates <= binary systems of two remnants
 - X ray binary systems with Bhs <= binary systems with one remnant
 - Non-interacting: Thompson et al. 2019 (mass-gap object)

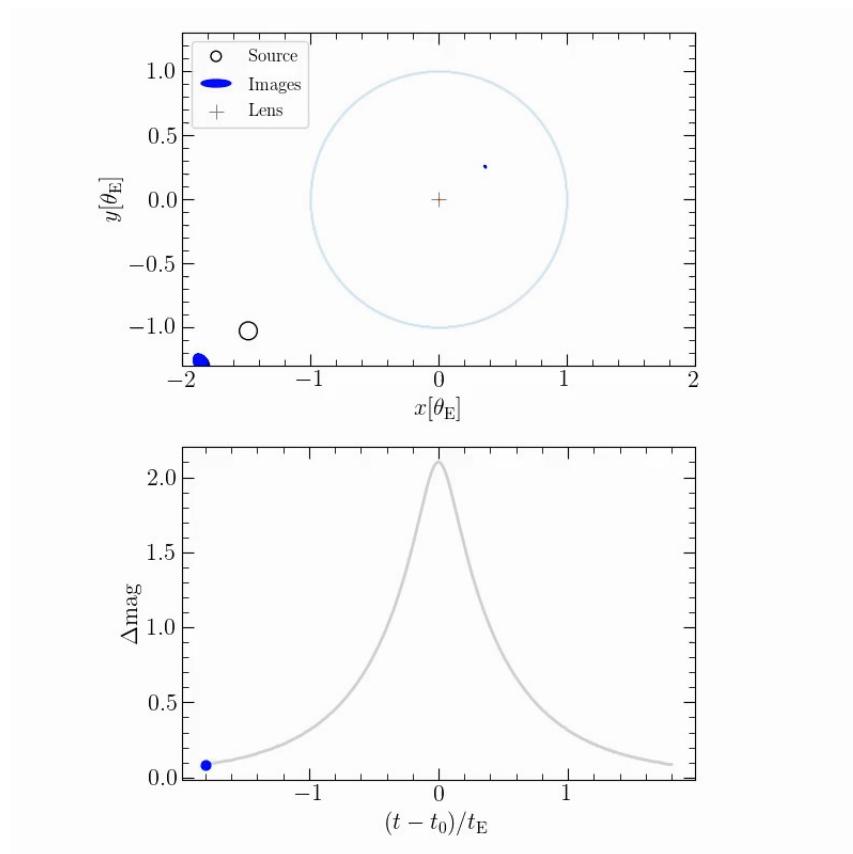
Gravitational Microlensing

- Change in brightness of a background source star due to gravitational field bending light by a foreground lens
- Can detect massive objects which don't emit light:
- Planets, white dwarfs, neutron stars, **black holes**
- Can be used to study the structure of the Milky Way (optical depth for microlensing)



Source: K.A. Rybicki

Gravitational Microlensing



Source: <https://www.astrouw.edu.pl/~krybicki/animations.php>

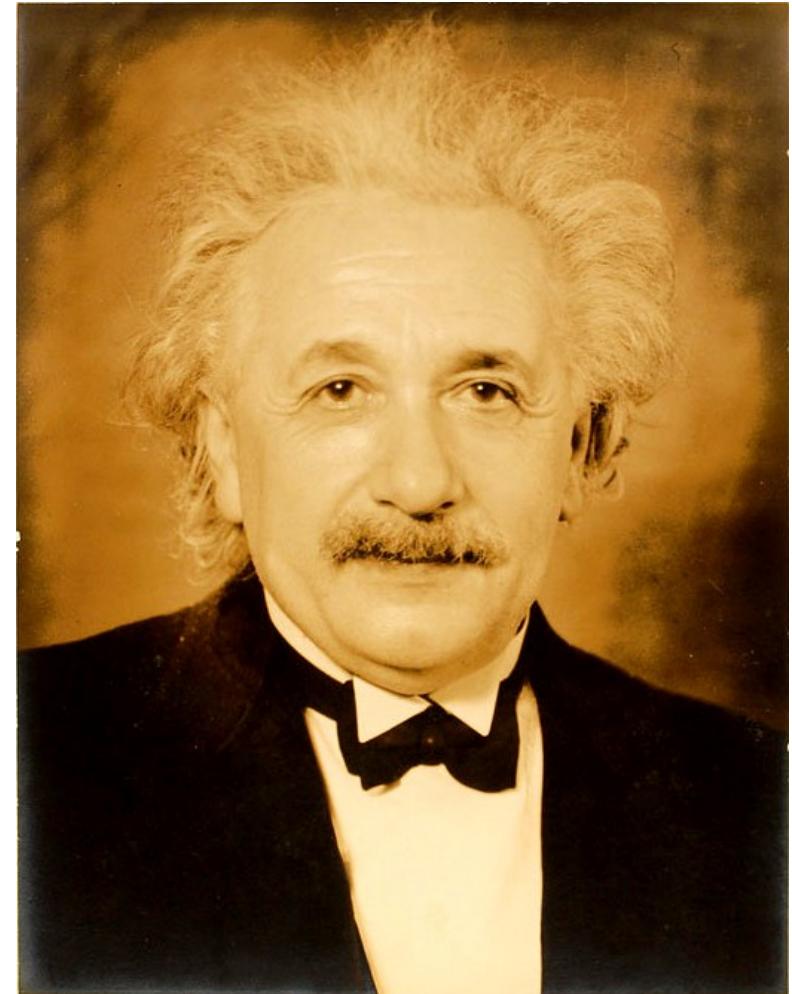
Gravitational Microlensing

- First by Einstein (1936),
by suggestion of R. W.
Mandl

[https://www.mpiwg-berlin.mpg.de/Preprints
/P160.PDF](https://www.mpiwg-berlin.mpg.de/Preprints/P160.PDF)

[https://www.sciencenews.org/blog/context/
amateur-who-helped-einstein-see-light](https://www.sciencenews.org/blog/context/amateur-who-helped-einstein-see-light)

- Theoretical ground-work:Liebes (1964), Refsdal (1964)
- Revival: Paczynski (1986)



Einstein in 1935, Princeton,
Source: https://en.wikipedia.org/wiki/Albert_Einstein

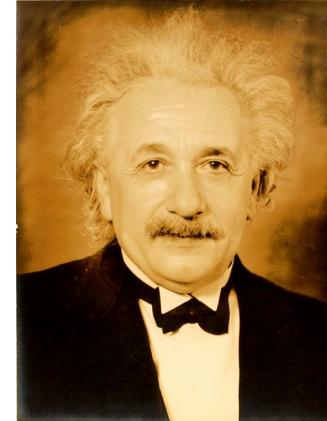
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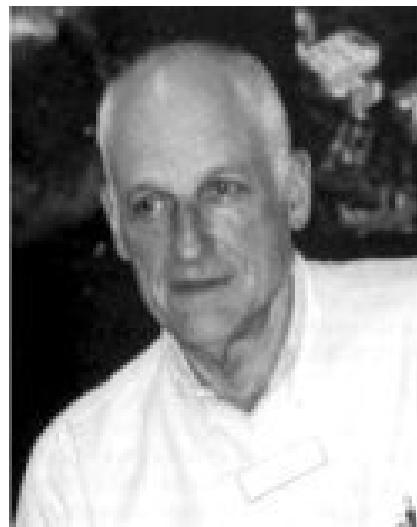
PROPOSES EINSTEIN TEST

Rudi W. Mandl, as he washed dishes in a restaurant, worked out an astronomical test of the relativity theory that won the approval of Professor Einstein and was published by him.

Source:
<https://www.sciencenews.org/archive/attic-genius-wihns-einstiens-approval-relativity-test>

Gravitational Microlensing

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Liebes (1964),
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Sidney Liebes



Sjur Refsdal

Source:
https://www.globalcommunity.org/wtt/walk_bio.shtml
https://en.wikipedia.org/wiki/Sjur_Refsdal

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Gravitational Lenses

Sidney Liebes, Jr.

Phys. Rev. 133, B835 – Published 10 February 1964

Article

References

Citing Articles (195)

PDF

Export Citation

ABSTRACT

A stellar gravitational lens has the capacity to intensify by a factor in excess of 1000 the image of

THE GRAVITATIONAL LENS EFFECT*

Sjur Refsdal

(Communicated by H. Bondi)

(Received 1964 January 27)

Summary

The so-called gravitational lens effect, previously worked out by Tikhov in 1937, is derived in a simple manner. The effect is caused by the gravitational deflection of light from a star *S* in the gravitational field of another star *B*, and occurs when *S* lies far behind *B*, but close to the line of sight through *B*. It turns out that a considerable increase in the apparent

Gravitational Microlensing

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- Theoretical ground-work:
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THE ASTROPHYSICAL JOURNAL, 304: 1–5, 1986 May 1
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GRAVITATIONAL MICROLENSING BY THE GALACTIC HALO

BOHDAN PACZYŃSKI¹

Princeton University Observatory

Received 1985 August 1; accepted 1985 October 23

ABSTRACT

Source:
<https://www.nature.com/articles/447106a>

Gravitational Microlensing

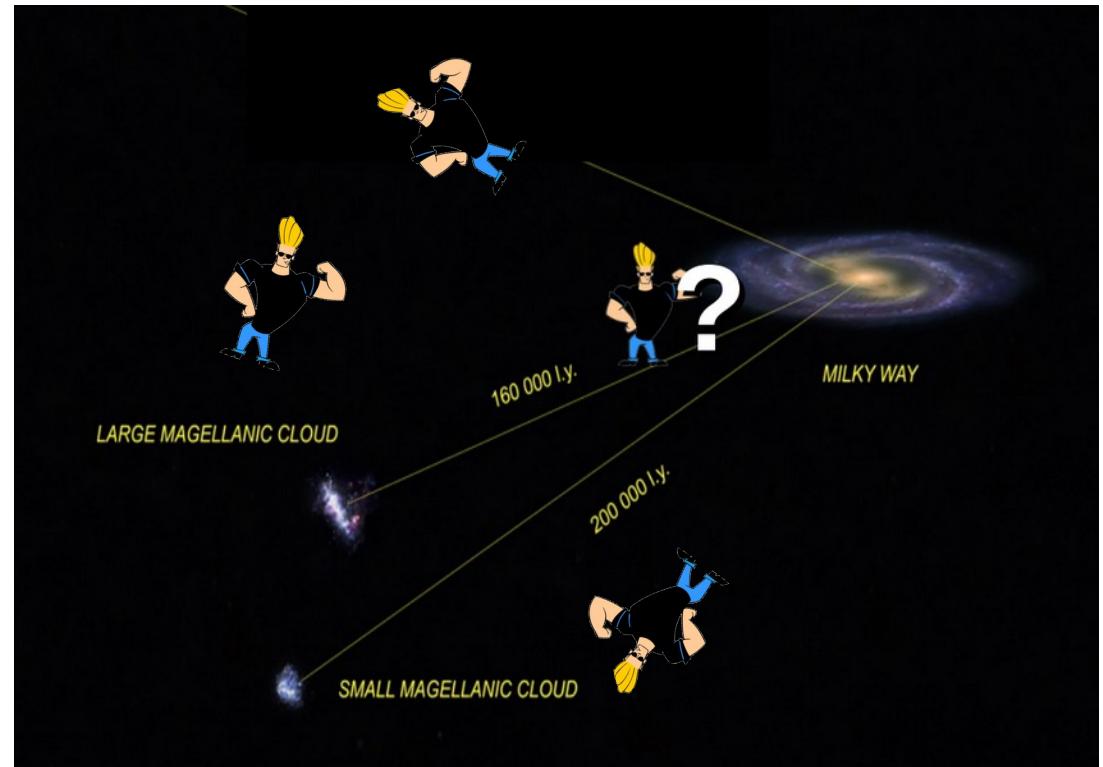
- Bohdan Paczyński → observing LMC (MACHOs)



WIMP
Weakly
Interacting
Massive
Particle



MACHO
Massive
Compact
Halo
Object

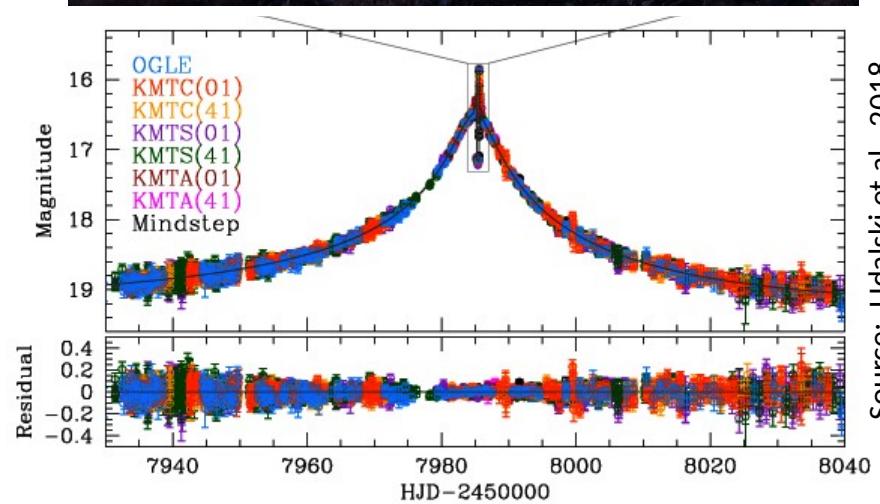


Gravitational Microlensing

- **Surveys:** MACHO, OGLE, EROS, MOA, KMTNet, ZTF
- **MACHOs** → initially proven to exist by MACHO survey (Alcock+ 1997, Bennett+ 2005), but soon ruled out by EROS (Tisserand+ 2007) and OGLE (Wyrzykowski+ 2009, 2010, 2011a,b)
- **Modern uses:**
 - Search for planets



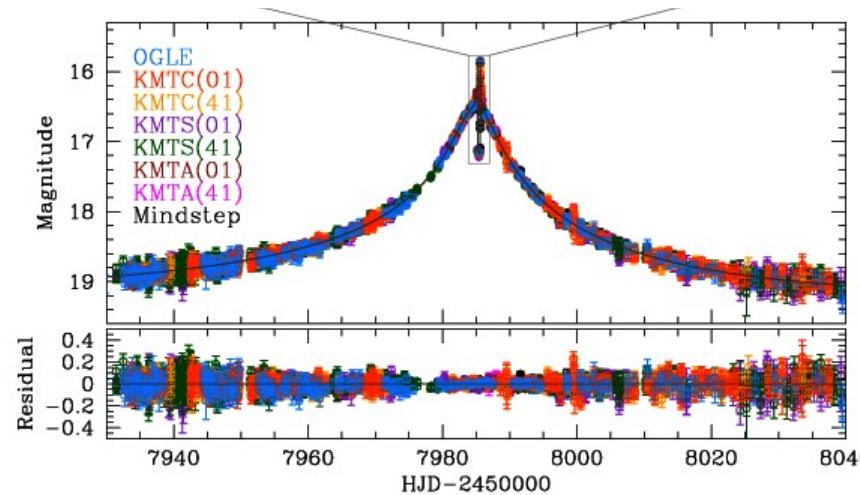
Source: OGLE/KU-tW



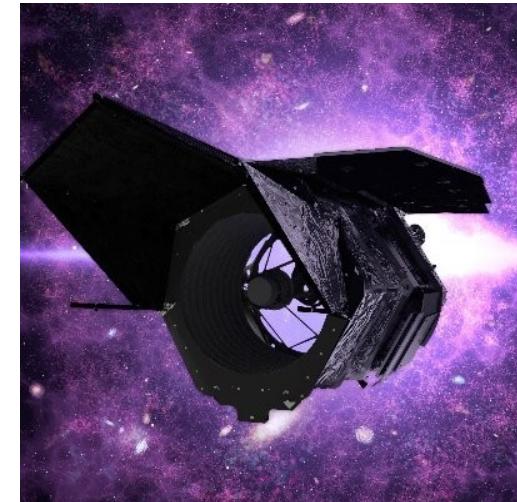
Source: Udalski et al., 2018, 1
Acta Astron., 68,

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- **Modern uses:**
 - Search for planets
 - Future: Nancy Grace Roman Space Telescope



Source: Udalski et al., 2018,
Acta Astron., 68, 1

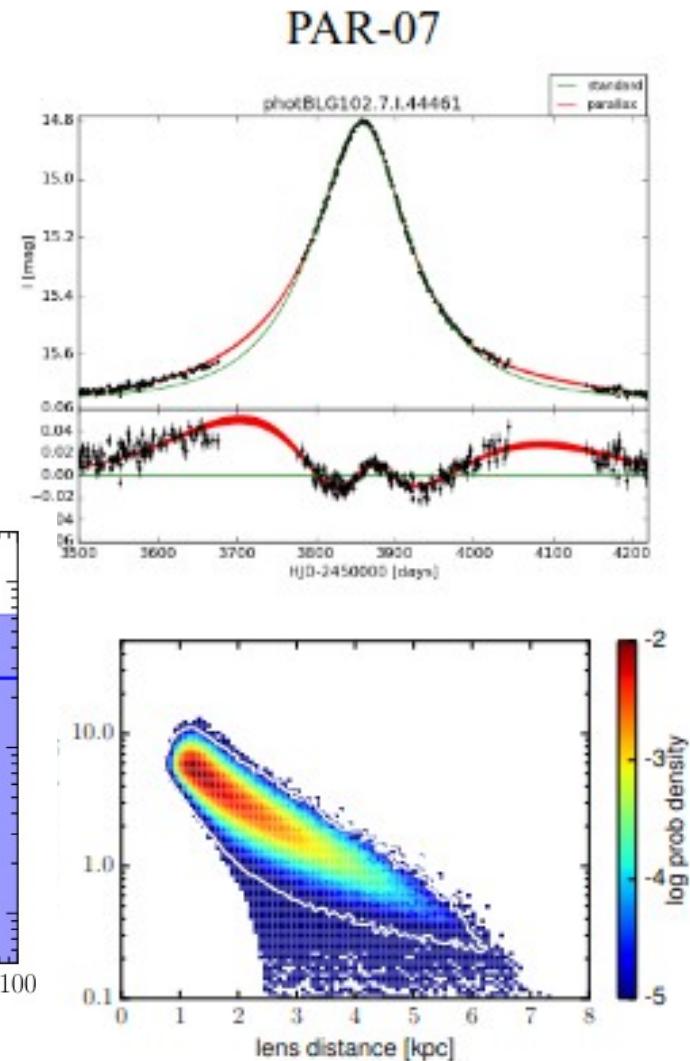
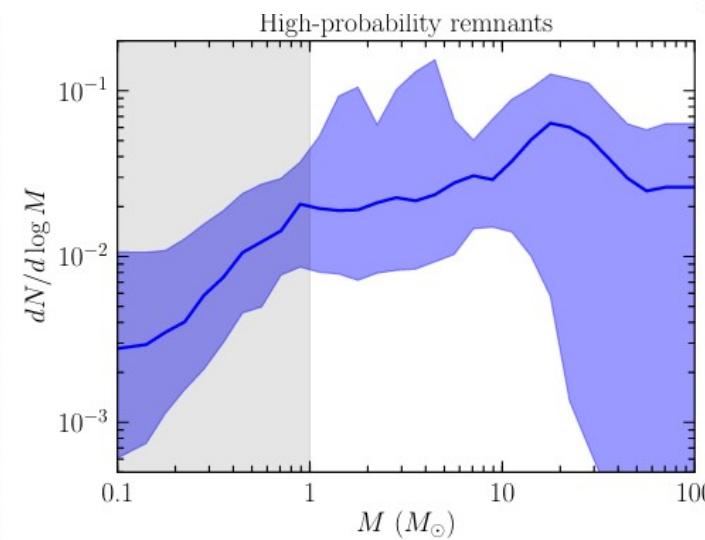
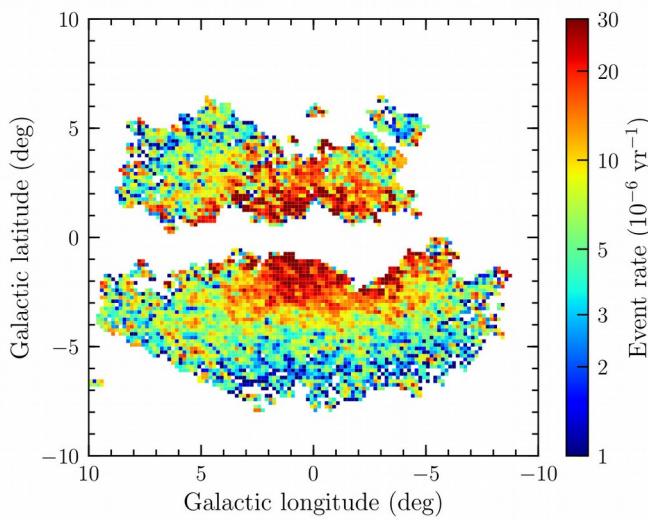


Source: NASA

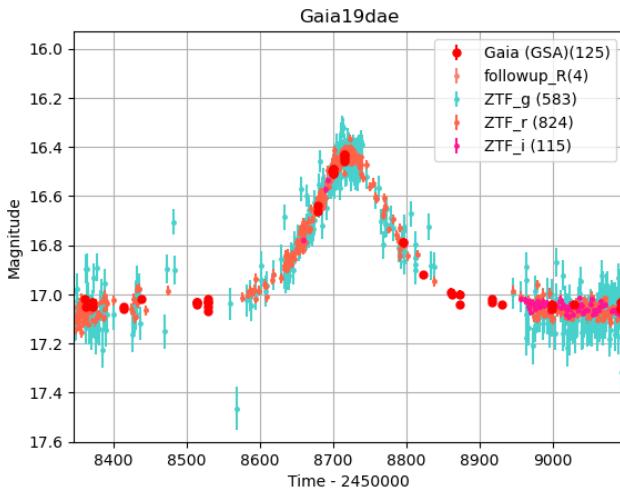
Gravitational Microlensing - history

- Other uses:

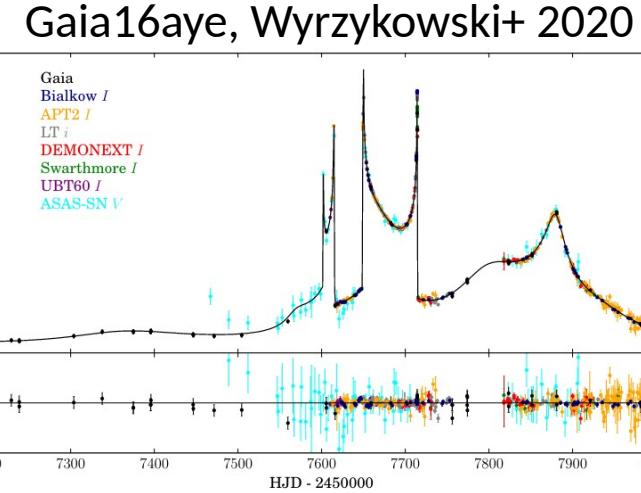
- Dark remnants (Wyrzykowski et al. 2016, Wyrzykowski&Mandel 2020, Mróz et al. 2021)
- Structure of the Galaxy (Wyrzykowski et al. 2015, Mróz et al. 2020a,b)



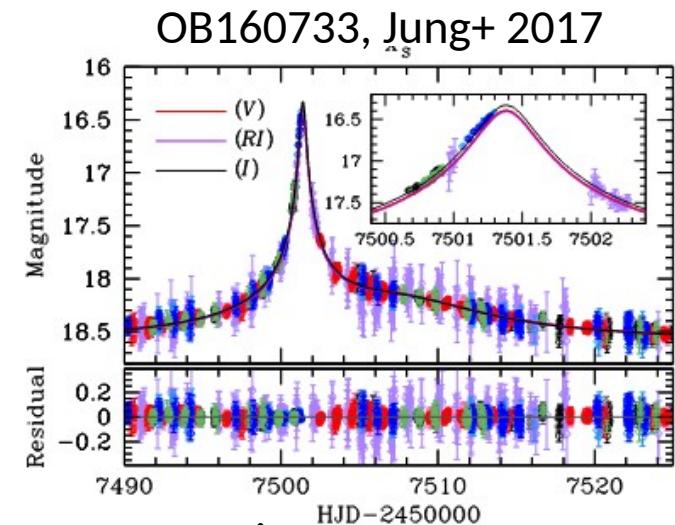
Microlensing events zoo



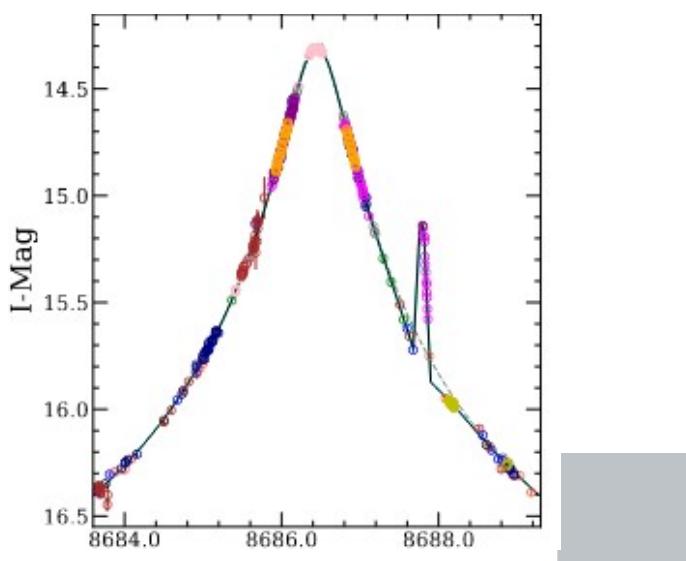
Single lens event



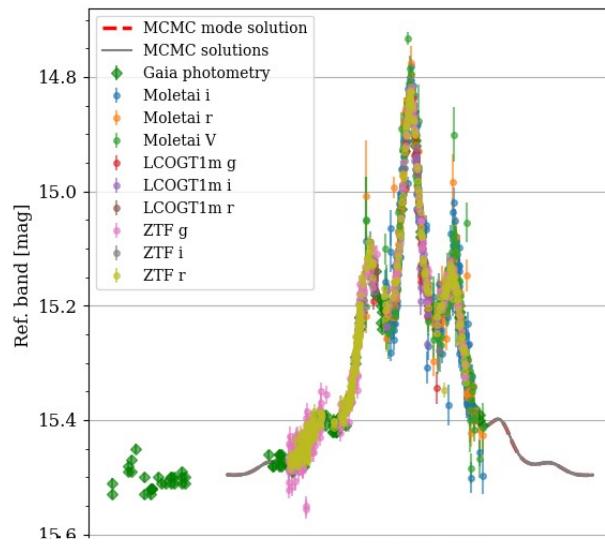
Binary lens event



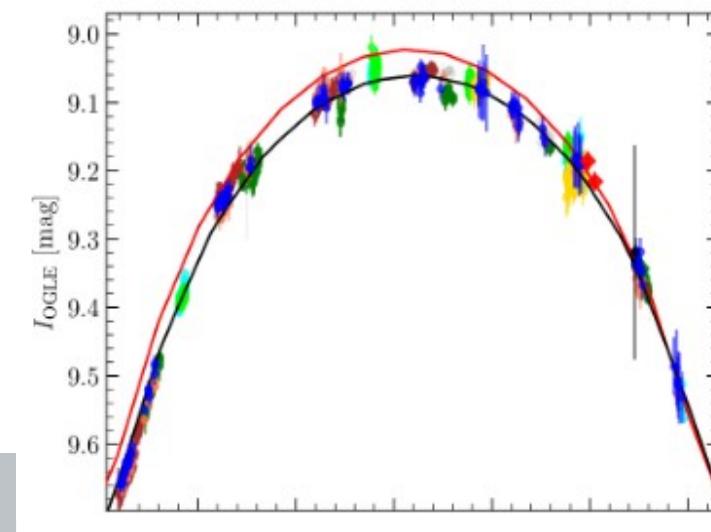
Binary source event



Planetary event



Gaia19dke, Maskoliunas+ in prep.



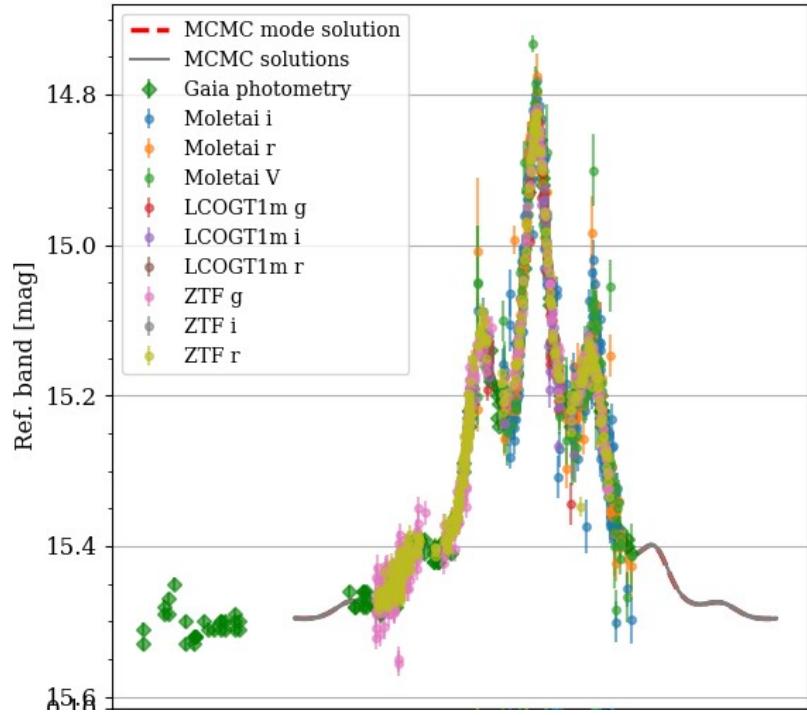
Parallax effect

- Movement of earth starts to be reflected in the light-curve
- For longer events → the lens might be massive (or moves slowly...)

$$M_L = \frac{\theta_E}{\kappa \pi_E}$$

θ_E is circled in red and points to the Einstein ring radius.

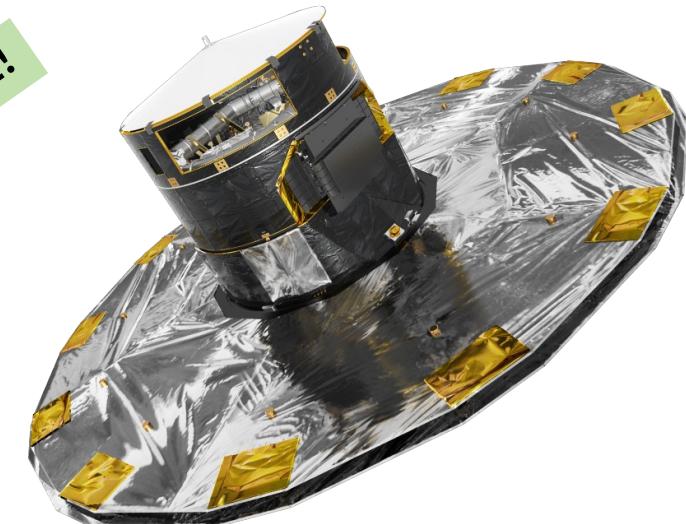
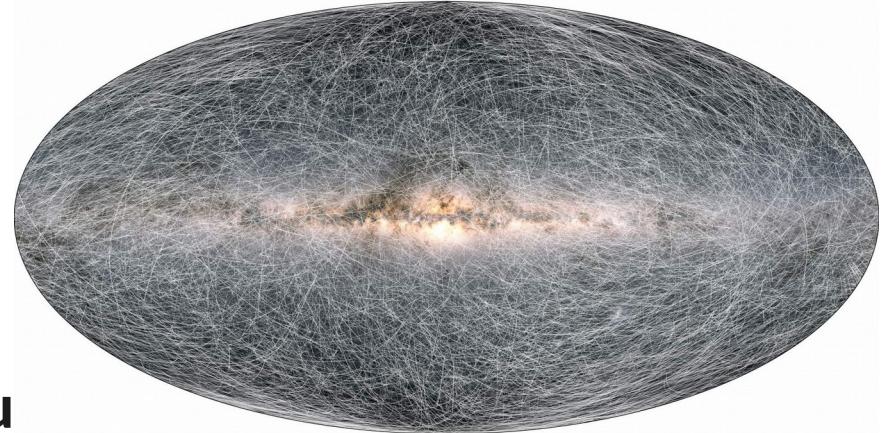
κπ_E is circled in red and points to the microlensing parallax.



Gaia19dke, Maskoliniunas+ in prep.

Gaia

- Launched on 19th Dec 2013 from French Guiana
- Space satellite in L2 point (distance = 1% au)
- Main goal: astrometry for 1 billion sou MW with precision down to 24 mas
- Whole sky covered 40-200 times over 5 years, cadence: 9x4s - 106min - 30 days DONE!
- Astrometric time-series!
- Data Release 2 => over 2000 papers
- Data Release 3 coming very soon!

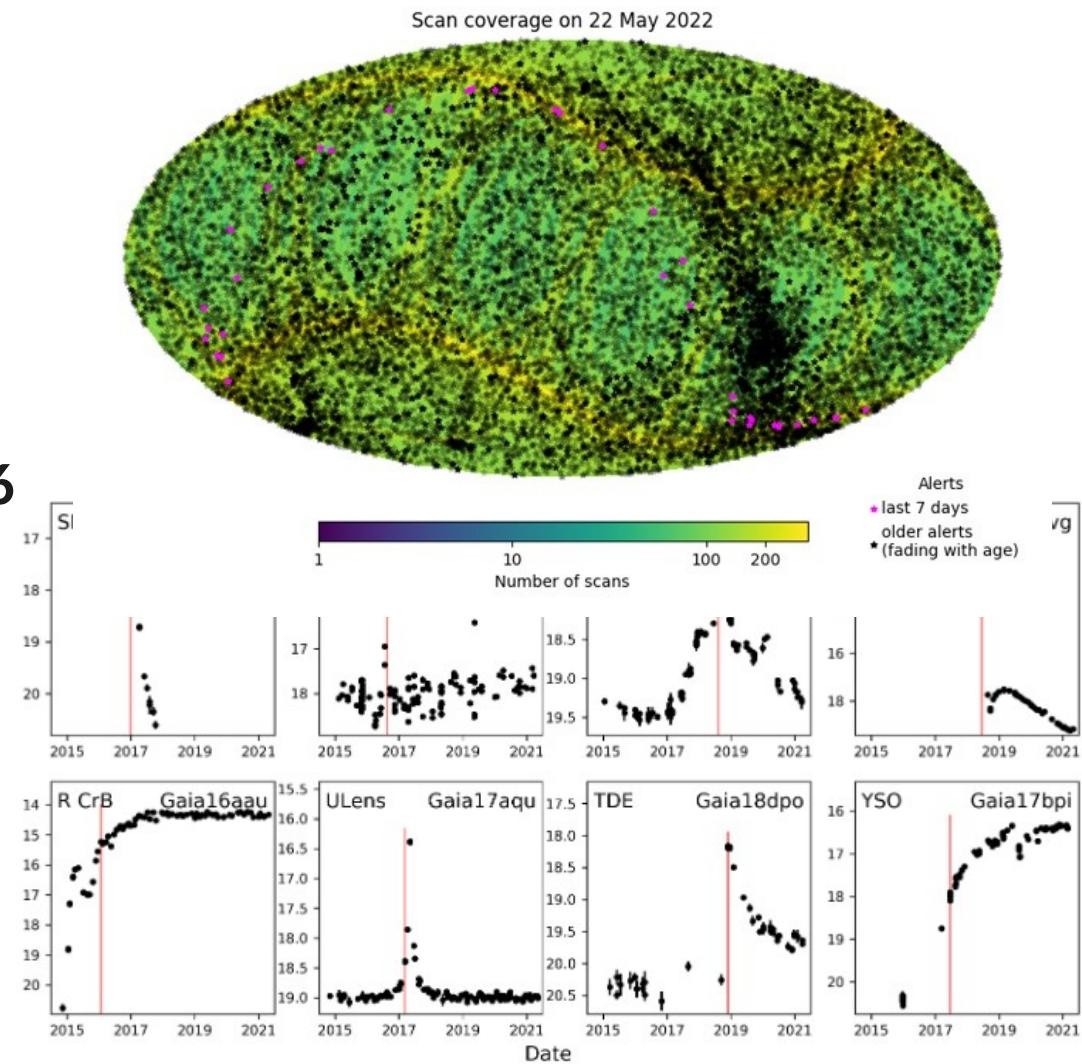
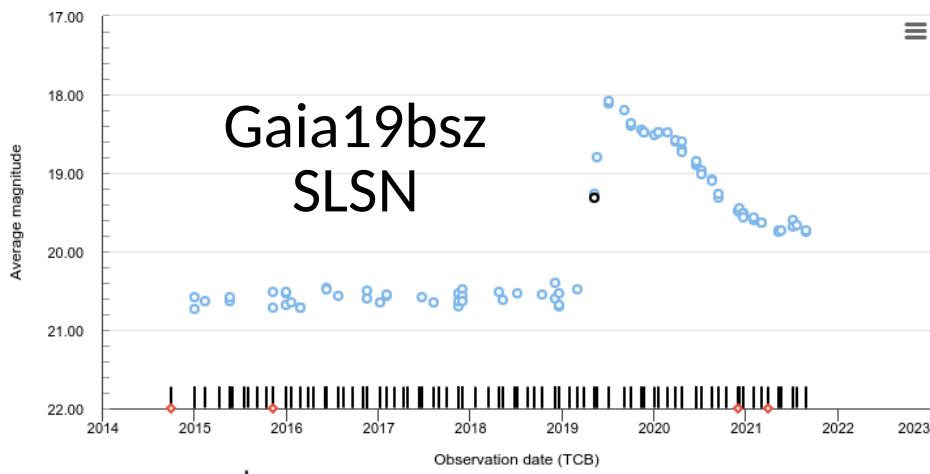


Source: ESA/Gaia/DPAC,
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Source: ESA/Gaia/DPAC

Gaia Science Alerts

- Publishing since 2015
 - 19 000+ Alerts on transients
 - Mostly SNs, novae, CVs, YSOs...
- <http://gsaweb.ast.cam.ac.uk/alerts>
- 500+ microlensing events since 2016



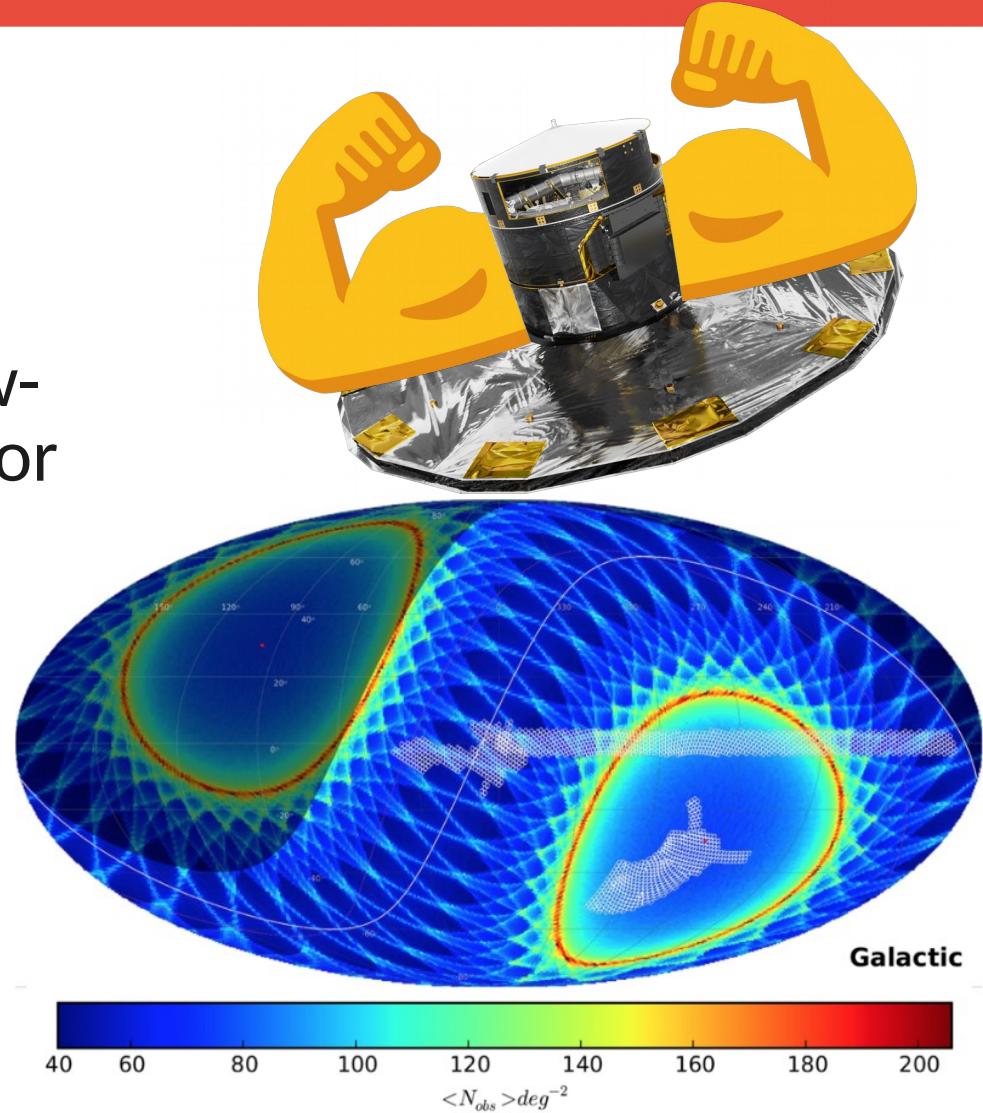
Gaia and microlensing

- **Strengths**

- Observes entire Galaxy
- Three band photometry, low-resolution spectra – useful for classification
- Astrometric time-series!

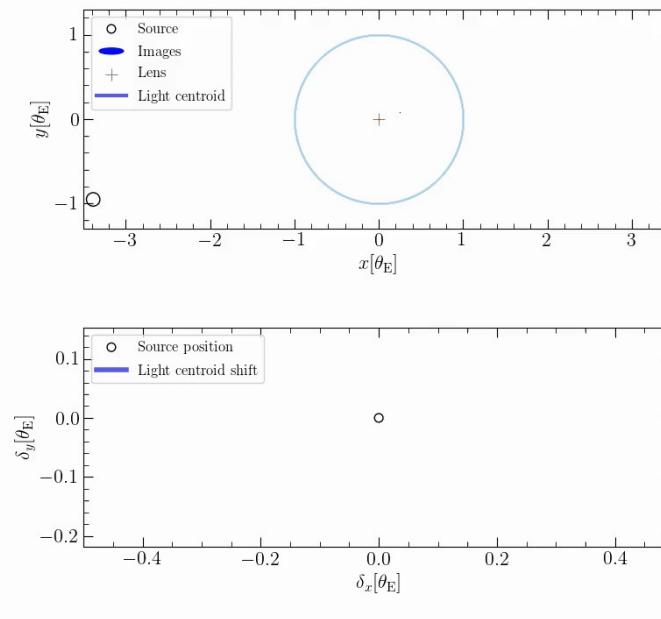
- **Weaknesses**

- Low-cadence (especially in Galactic Bulge)



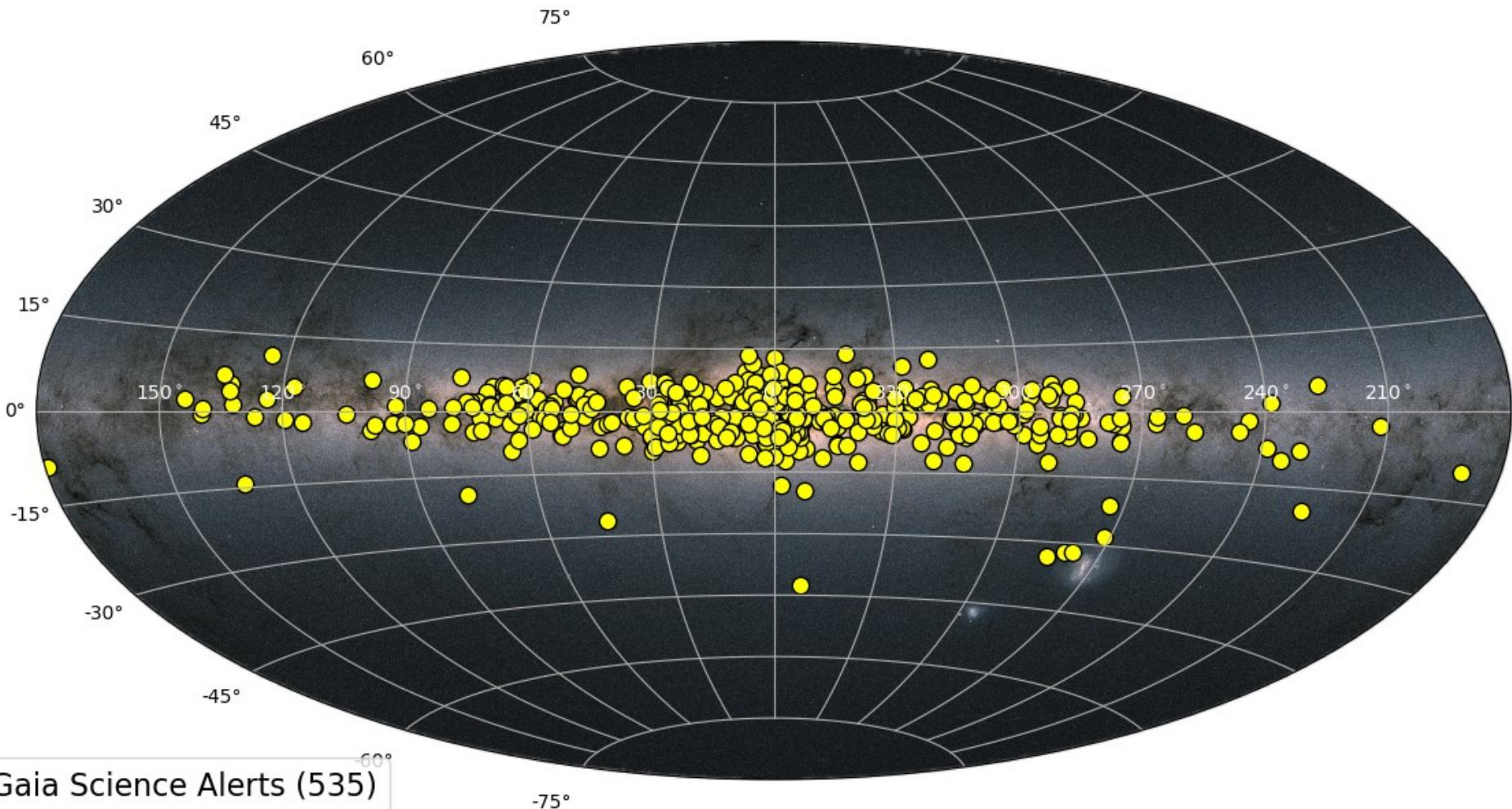
Gaia and microlensing

- Astrometric and photometric microlensing



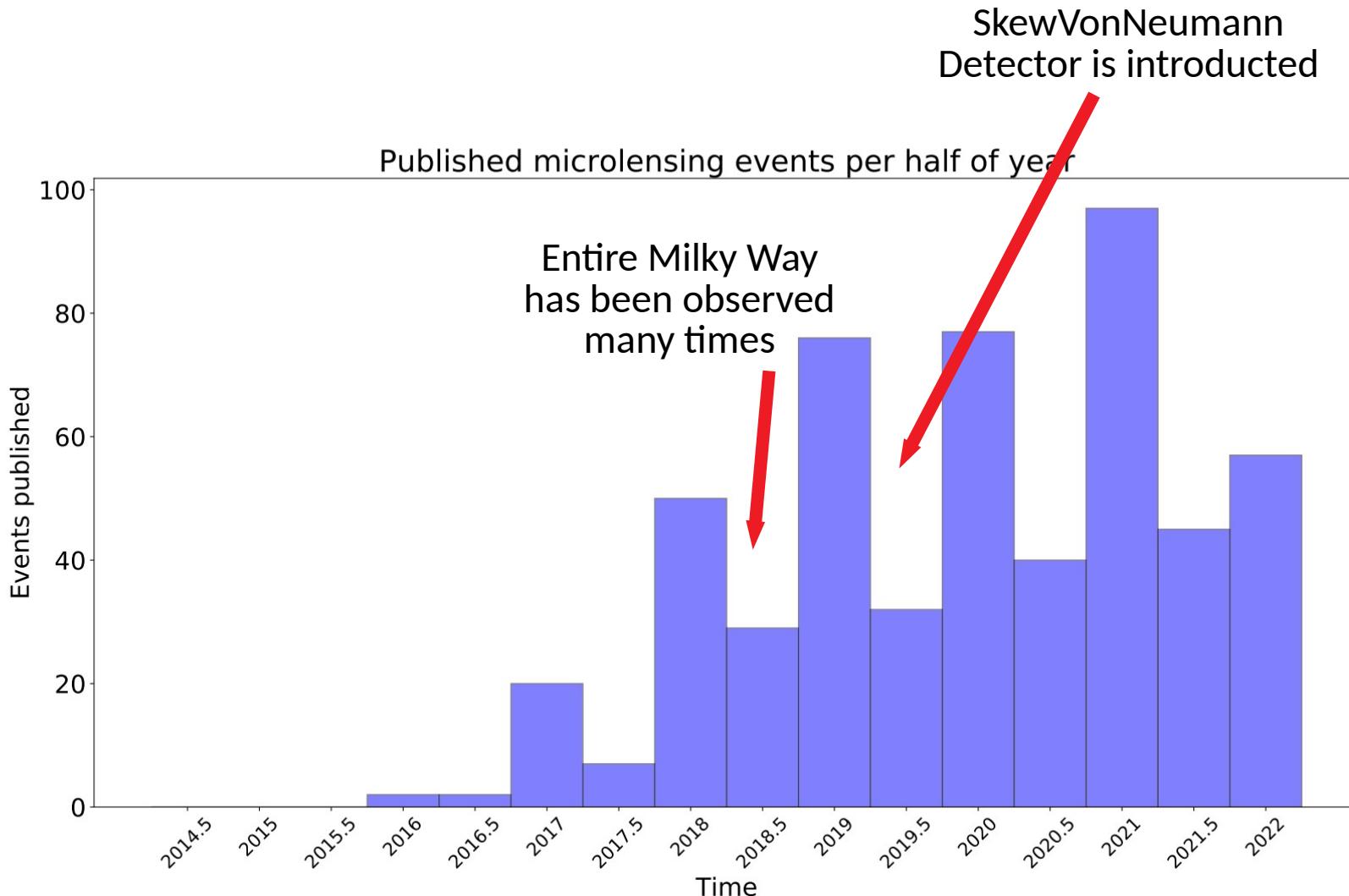
Source: <https://www.astrouw.edu.pl/~krybicki/animations.php>

Microlensing in Gaia Science Alerts



Background image: ESA/DPAC

Microlensing in Gaia Science Alerts over the years



How to get better data?

- Follow-up network! → Black Hole TOM
- ... and applying for time with various observatories (LCO, REM, SMARTS...)

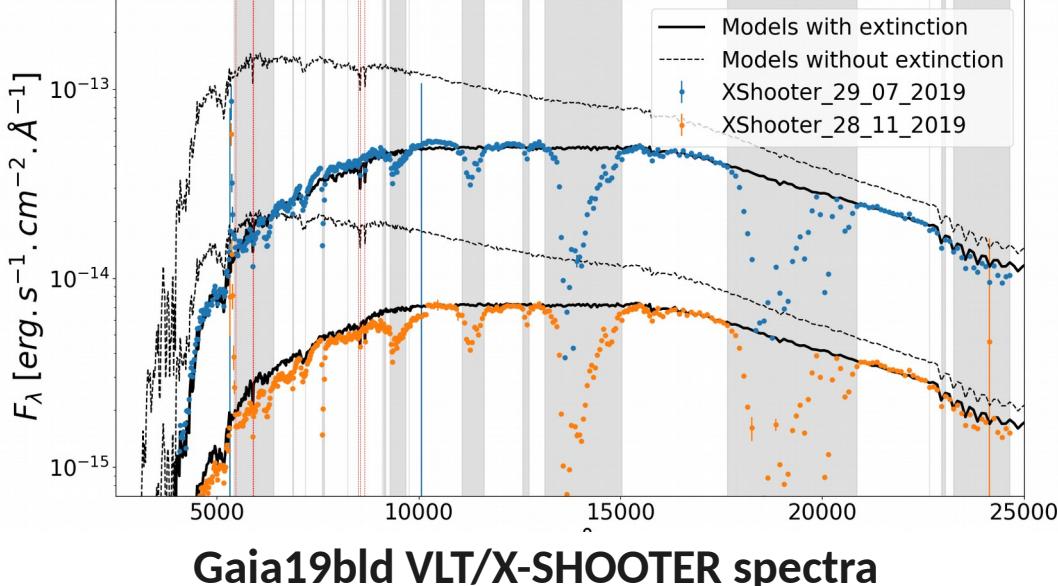
Follow-up network



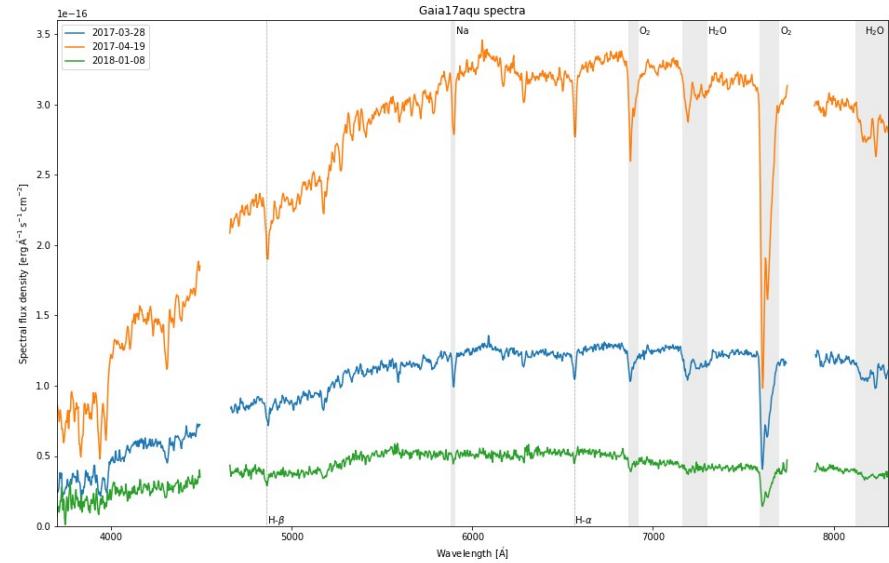
How to get better data?

- Follow-up network! → Black Hole TOM
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- Photometry and spectroscopy
 - Low res spectroscopy on LT/SPRAT... → classification
 - High res spectroscopy on RSS/SALT, VLT/XSHOOTER → source star parameters determination

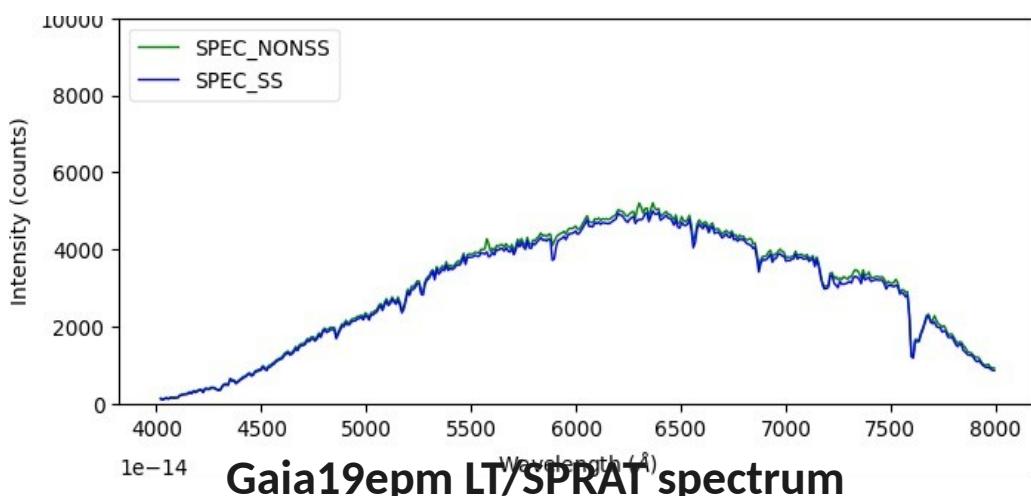
Highlights



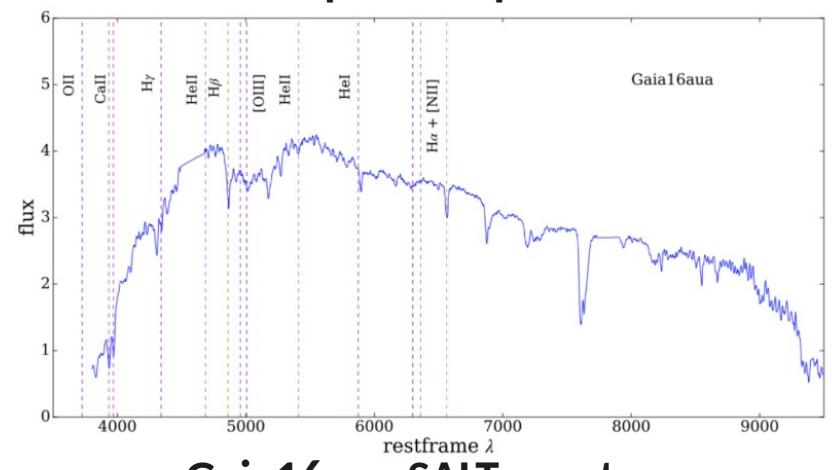
Gaia19bld VLT/X-SHOOTER spectra



Gaia17aqu SALT spectra



Gaia19epm LT/SPRAT spectrum

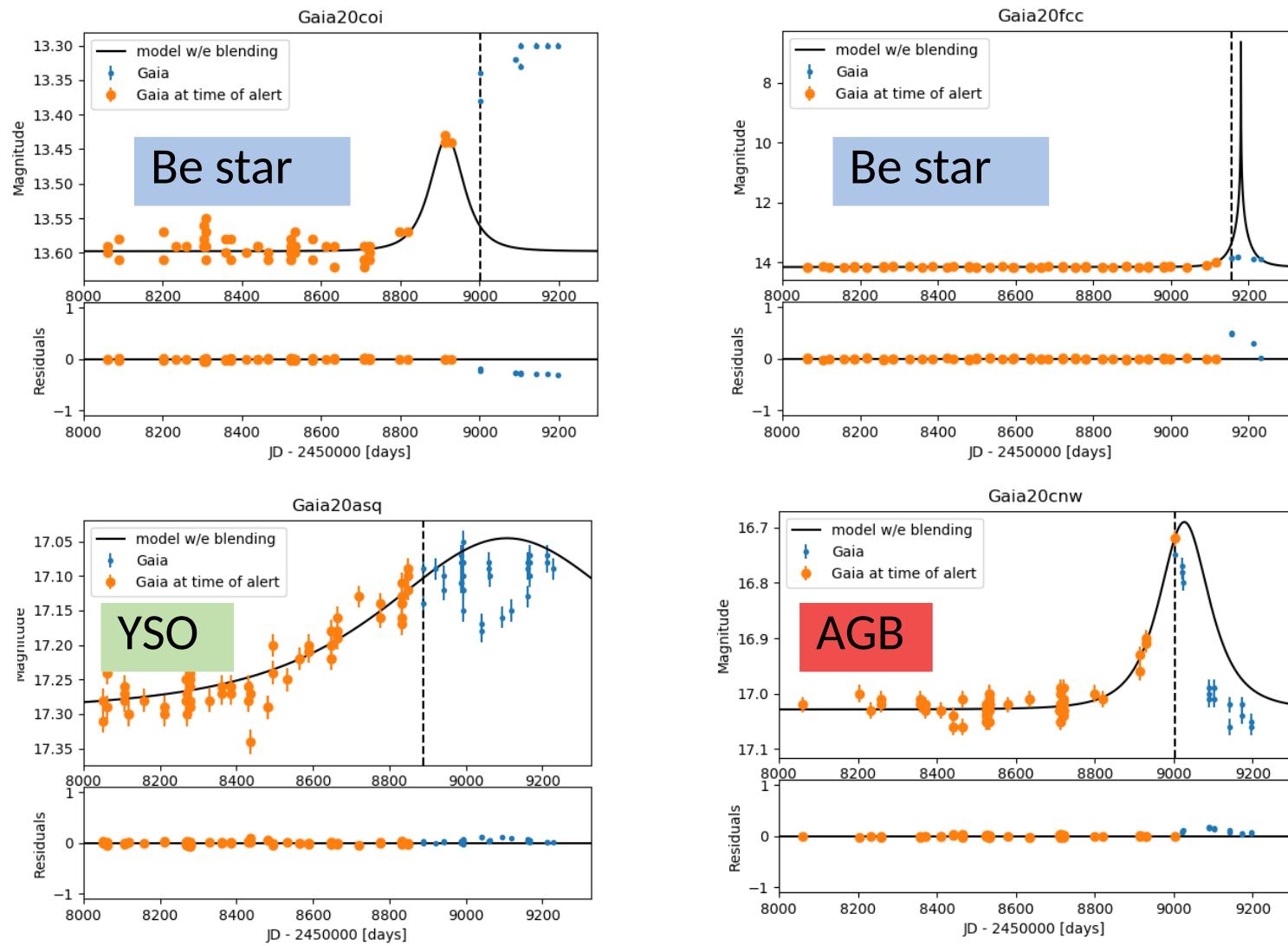


Gaia16aua SALT spectrum

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- Early classification based on archival data

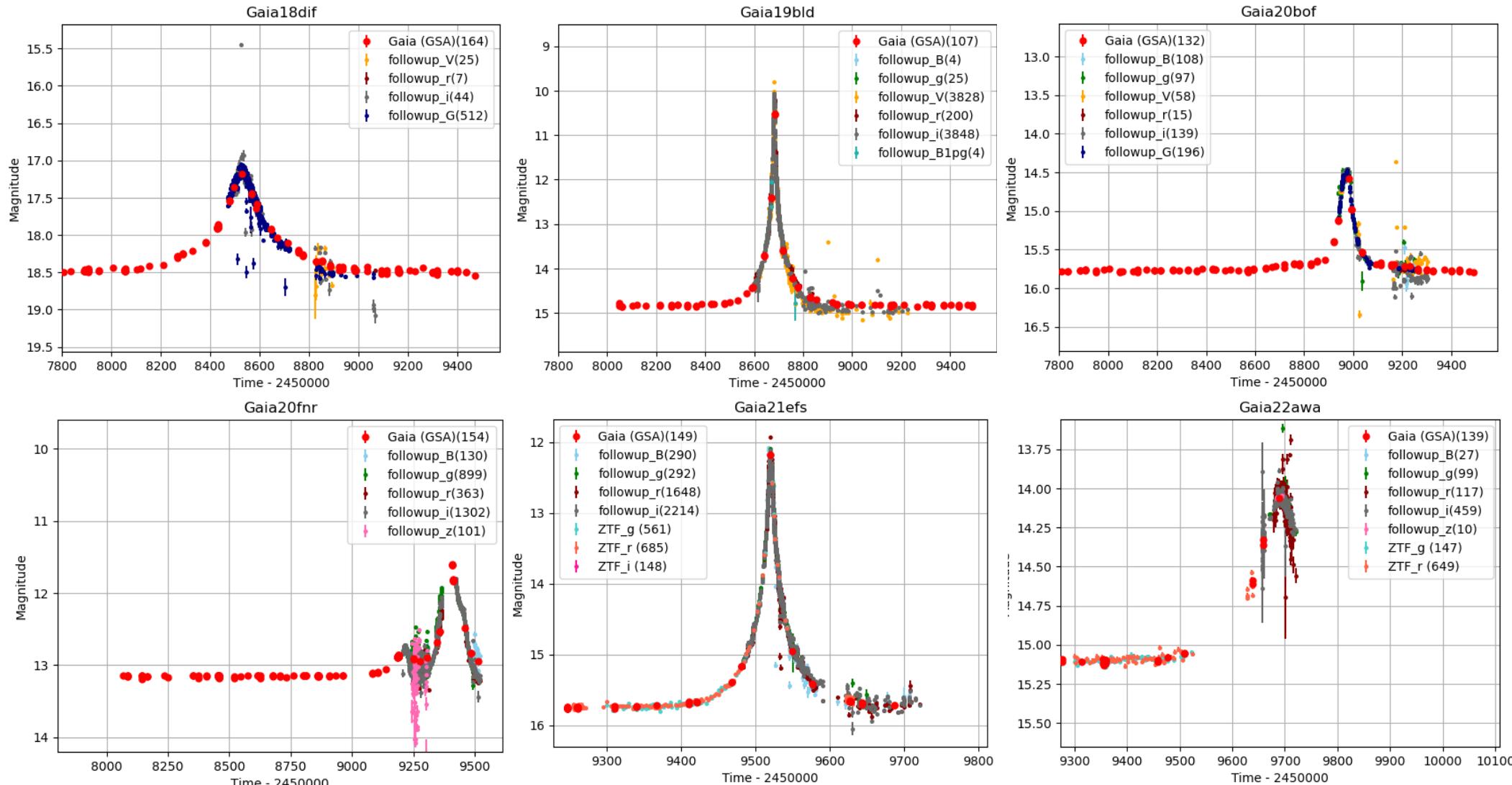
Most Common Contaminants



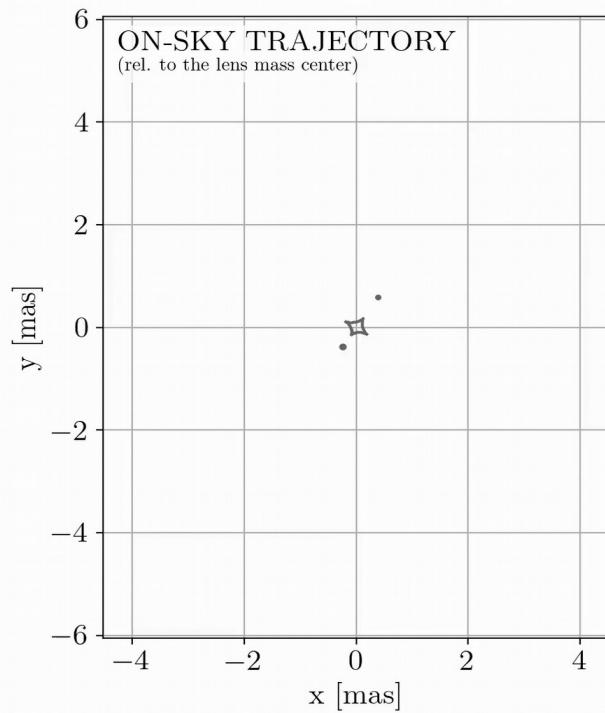
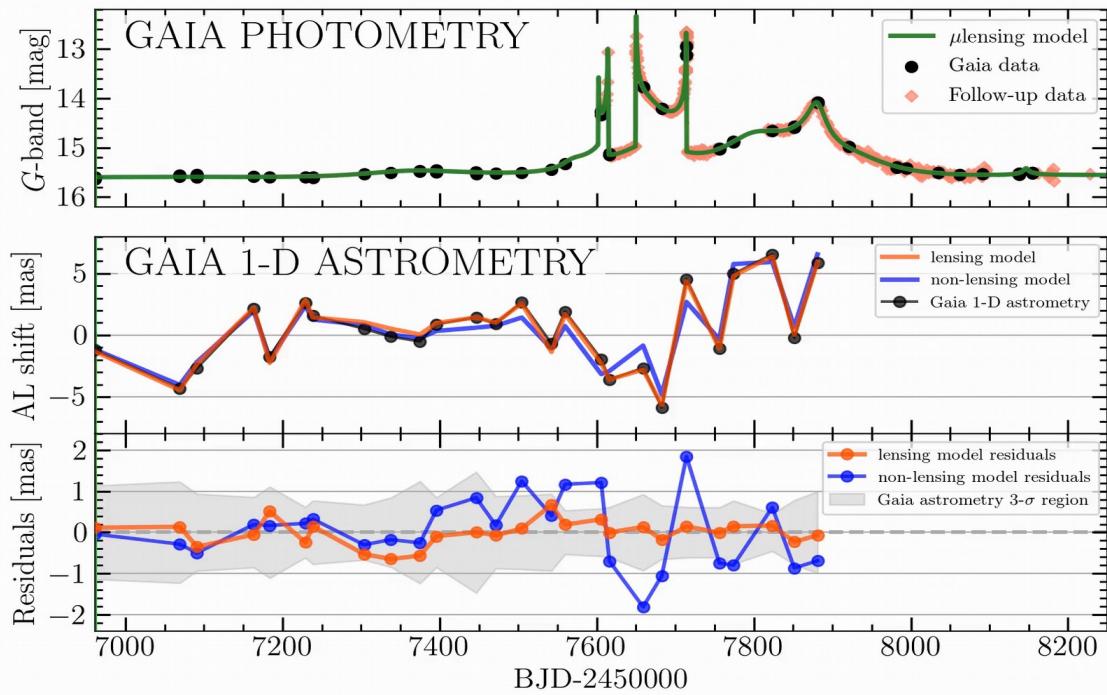
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- Photometry and spectroscopy
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- Early classification based on archival data
- The outcomes → catching interesting events!

Microlensing highlights - with follow-up



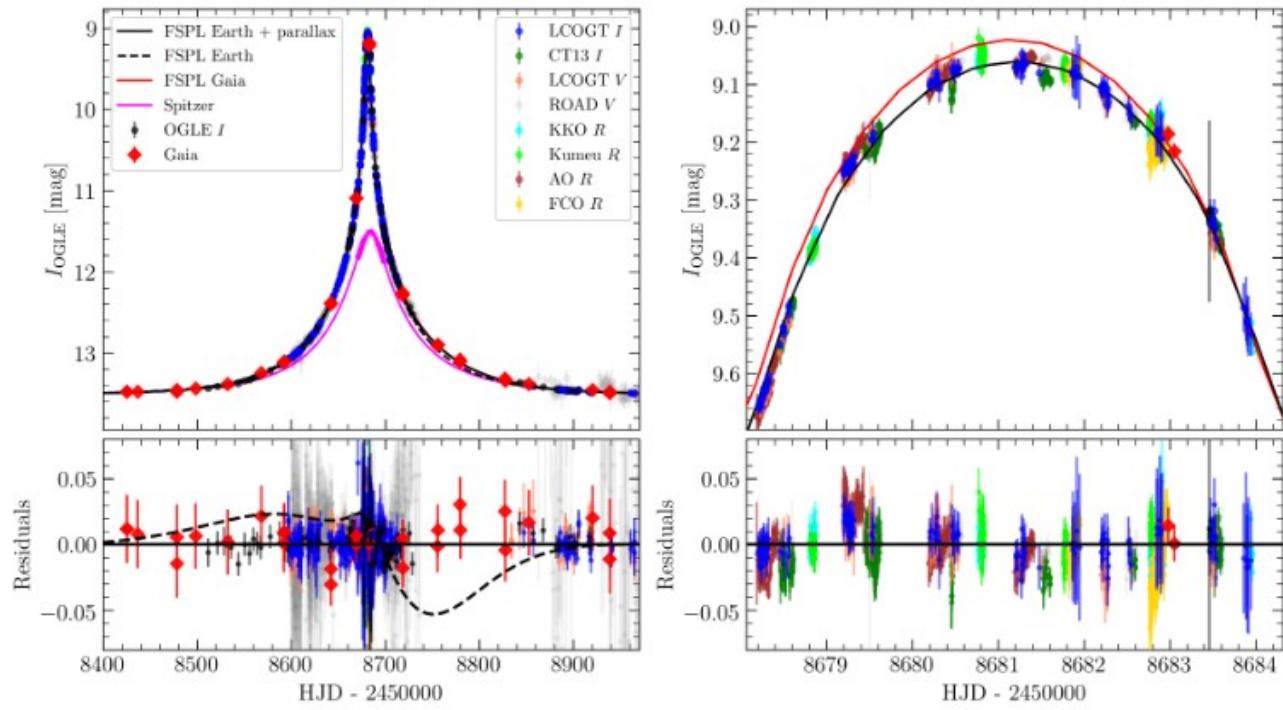
Gaia16aye



Source: <https://www.astrouw.edu.pl/~krybicki/animations.php>
https://www.cosmos.esa.int/web/gaia/iow_20210924

Gaia19bld

- **Microlensing event in the Galactic disc**
- **Observed (among others) by Spitzer and VLTI/PIONIER**
- **Also spectroscopic follow-up**

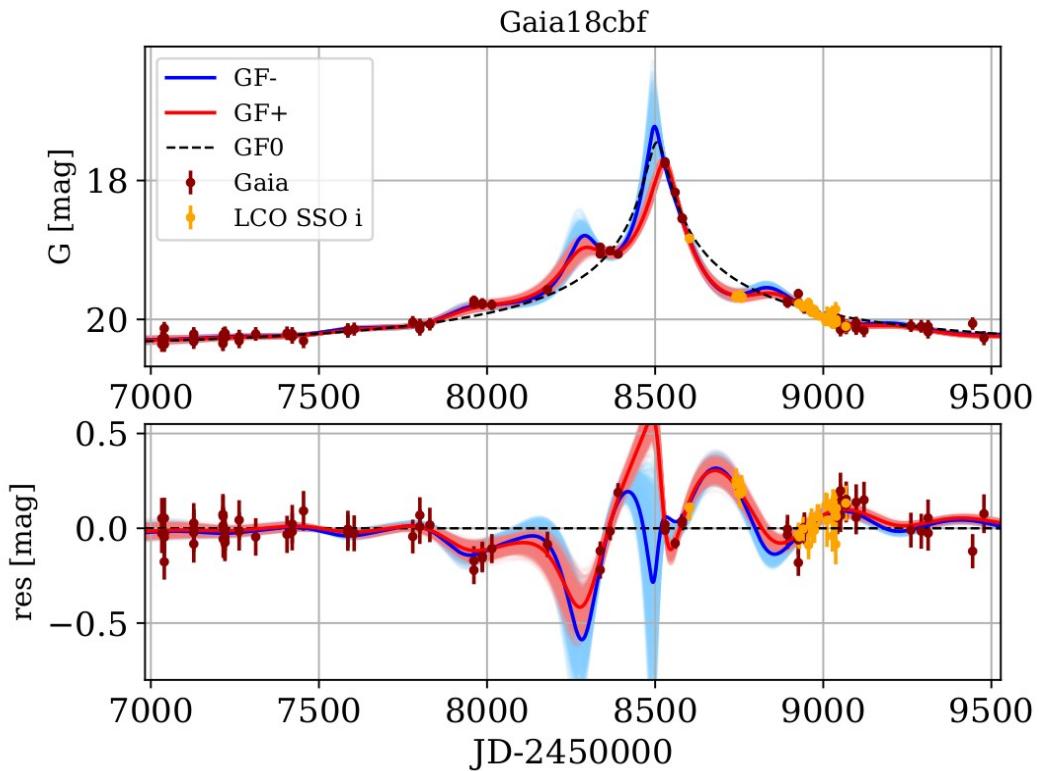


Animation: <https://www.astrouw.edu.pl/~krybicki/animations.php>

Gaia18cbf

- 2nd longest microlensing event ever observed in terms of t_E
- time span of the whole event of ~ 1900 days
- Very dim (baseline at $G \sim 20.5$ mag)

Parameter	GF0	GF+	GF-
$t_{0,\text{par}} - 2450000$. [days]	—	8529.00	
$t_0 - 2450000$. [days]	$8505.34^{+2.04}_{-2.00}$	$8524.76^{+4.33}_{-4.57}$	$8513.14^{+4.68}_{-3487}$
u_0	$0.0133^{+0.0086}_{-0.0065}$	$0.0825^{+0.0231}_{-0.0208}$	$-0.0567^{+0.0162}_{-0.0169}$
t_E [days]	$2058.40^{+1956.82}_{-781.12}$	$491.41^{+128.31}_{-84.94}$	$453.74^{+178.69}_{-105.74}$
π_{EN}	—	$-0.1192^{+0.0273}_{-0.0285}$	$-0.1697^{+0.0516}_{-0.0577}$
π_{EE}	—	$-0.0442^{+0.0077}_{-0.0084}$	$-0.0257^{+0.0054}_{-0.0057}$
G_0 [mag]	$20.44^{+0.07}_{-0.07}$	$20.34^{+0.03}_{-0.03}$	$20.34^{+0.04}_{-0.03}$
$f_{S,G}$	$0.198^{+0.109}_{-0.090}$	$0.894^{+0.263}_{-0.226}$	$0.717^{+0.294}_{-0.233}$
i_0 [mag]	$20.06^{+0.07}_{-0.07}$	$20.01^{+0.06}_{-0.05}$	$20.00^{+0.07}_{-0.06}$
$f_{S,i}$	$0.157^{+0.084}_{-0.072}$	$0.996^{+0.250}_{-0.219}$	$0.746^{+0.253}_{-0.208}$
χ^2	143.98	40.87	46.69
$\frac{\chi^2}{dof}$	1.40	0.40	0.46



- Two solutions:
 - GF+ $t_E = 491.41$ days, $f_{s,G} = 0.894$
 - GF- $t_E = 453.74$ days, $f_{s,G} = 0.717$
- Follow-up data from LCO network (2m at Siding Springs Observatory)

Gaia18cbf

How can we constrain mass and rule out star scenario?

Microlensing parallax

+

Proper motions from Gaia

+

Estimate of the distance to the source

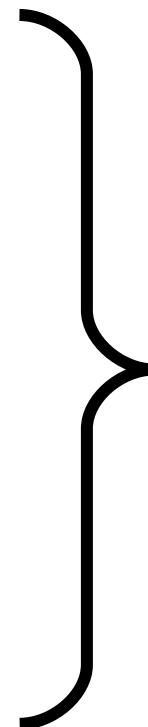
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Galactic model

+

Mass-function of remnants

Method described in: Wyrzykowski+ 2016, Wyrzykowski&Mandl 2020,
Mróz&Wyrzykowski 2021...



???



Gaia18cbf

How can we constrain mass and rule out star scenario?

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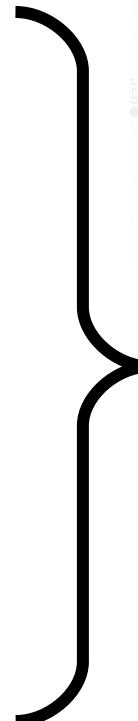
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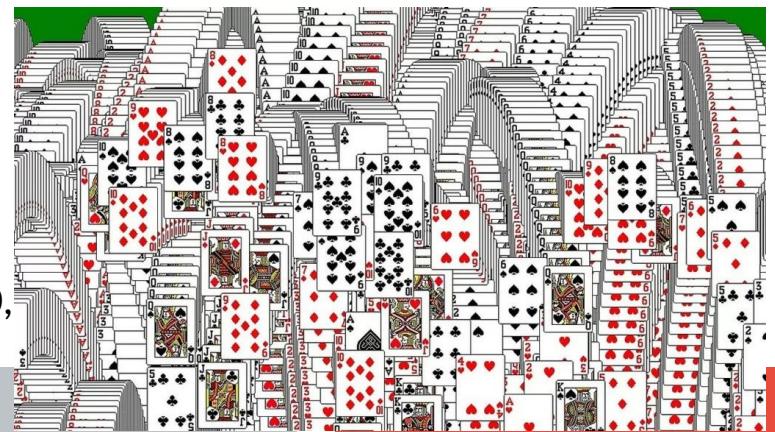
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Estimate of the distance to the source

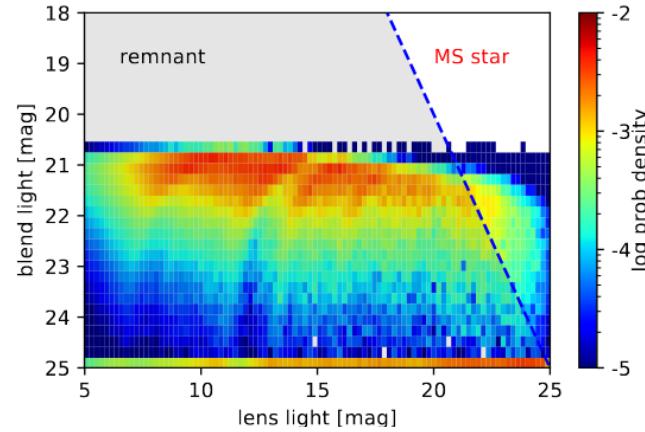
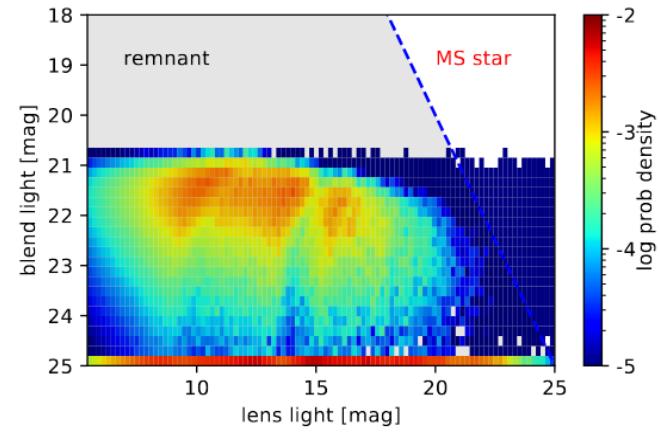
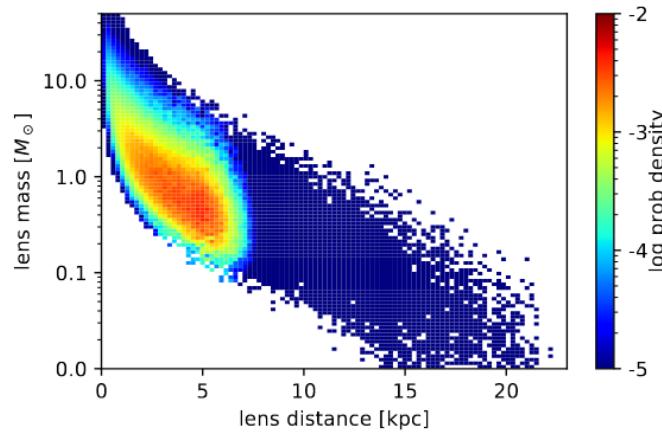
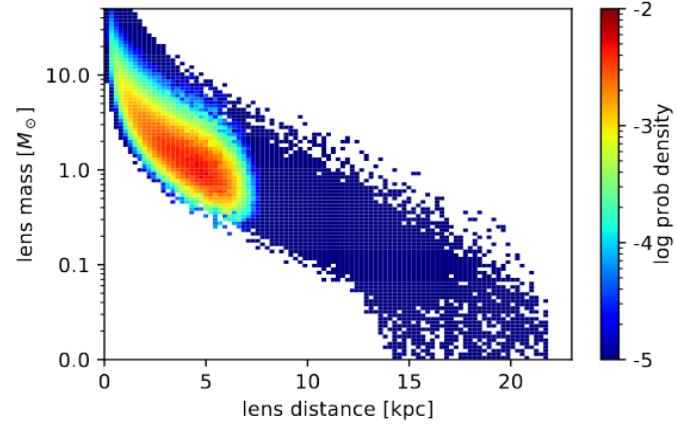
Method described in: Wyrzykowski+ 2016, Wyrzykowski&Mandl 2020,
Mróz&Wyrzykowski 2021... and still being refined



PROFIT



Gaia18cbf



GF+ lens is	$M_L [M_{\odot}]$	$2.65^{+5.09}_{-1.48}$
	$D_L [\text{kpc}]$	$2.84^{+1.94}_{-1.67}$

GF- lens is	$M_L [M_{\odot}]$	$1.71^{+3.78}_{-1.06}$
	$D_L [\text{kpc}]$	$2.66^{+1.97}_{-1.56}$

Gaia18cbf – different scenarios

- There was a number of assumptions done – let's play around with them:
 - Remnant mass function: could be less steep... →
 - or more steep ↘
 - Assumed distance to the source derived using archival photometric data...
 - but maybe Bailer-Jones distances are correct?
(Despite negative parallax) →

Parameter	GF+	GF-
G_{blend} [mag]	$21.31^{+\inf}_{-0.47}$	$20.89^{+\inf}_{-0.23}$
Mass function $\propto M^{-1.00}$, $14.5 \text{ kpc} < D_S < 23.0 \text{ kpc}$		
M_L [M_\odot]	$6.17^{+14.79}_{-4.09}$	$4.83^{+13.30}_{-3.41}$
D_L [kpc]	$1.55^{+2.08}_{-1.02}$	$1.45^{+2.00}_{-0.97}$
G_{MS} [mag]	<10.76 (B2.5V)	10.95 (B5V)
Mass function $\propto M^{-3.00}$, $14.5 \text{ kpc} < D_S < 23.0 \text{ kpc}$		
M_L [M_\odot]	$1.23^{+1.32}_{-0.53}$	$0.62^{+0.80}_{-0.30}$
D_L [kpc]	$4.34^{+1.27}_{-1.74}$	$4.14^{+1.35}_{-1.67}$
G_{MS} [mag]	17.81 (F6V)	21.89 (K8V)
Mass function $\propto M^{-1.75}$, $6.4 \text{ kpc} < D_S < 9.8 \text{ kpc}$ (B-J distance)		
M_L [M_\odot]	$3.78^{+7.78}_{-2.41}$	$2.38^{+5.73}_{-1.60}$
D_L [kpc]	$1.76^{+1.51}_{-1.01}$	$1.67^{+1.47}_{-0.97}$
G_{MS} [mag]	-0.39 (B7V)	13.13 (A0V)

<https://arxiv.org/abs/2111.08337>

The future

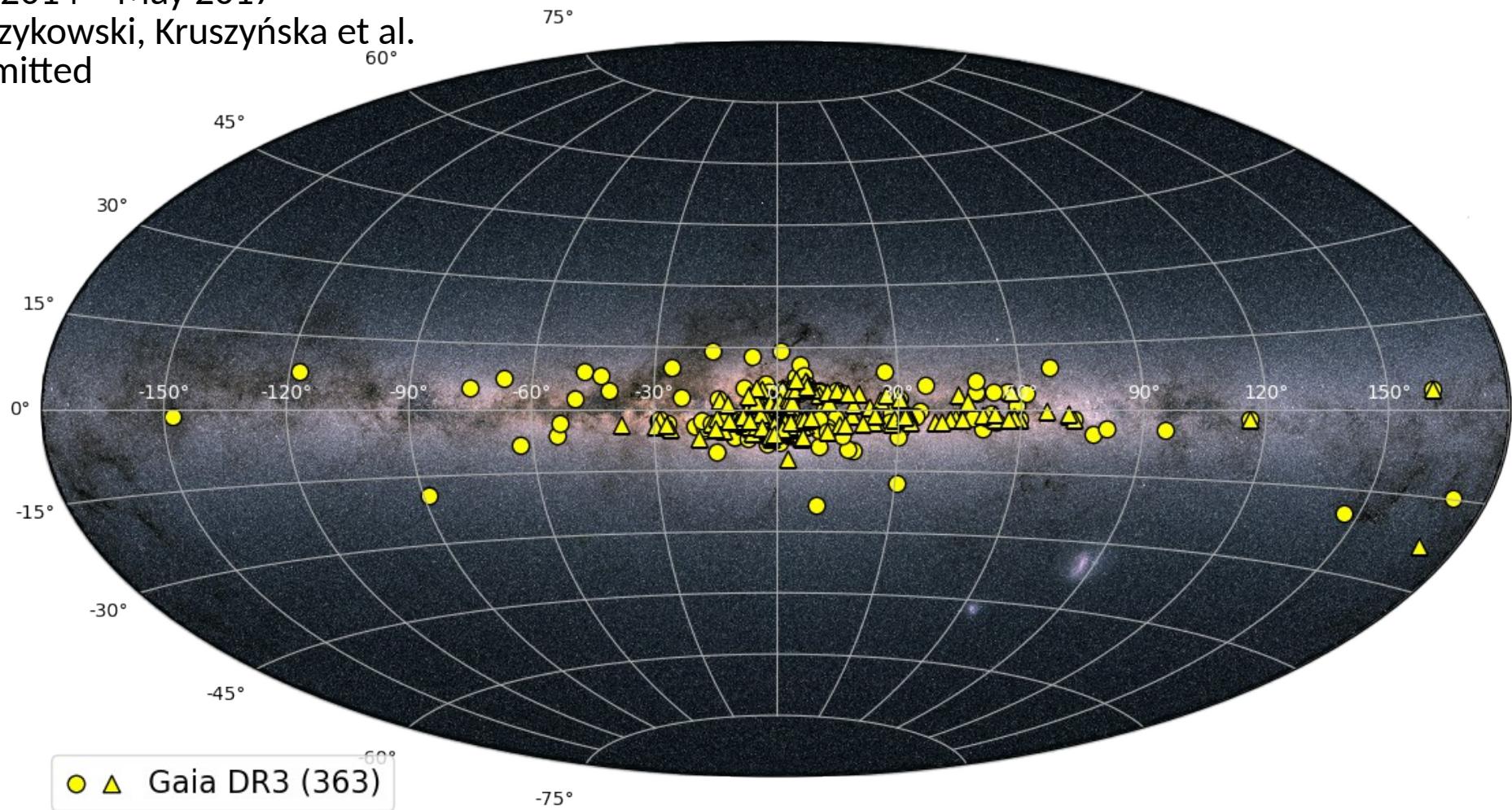
- Gaia DR3 microlensing catalogue

Gaia Data Release 3 - 13th June 2022

Candidate events from

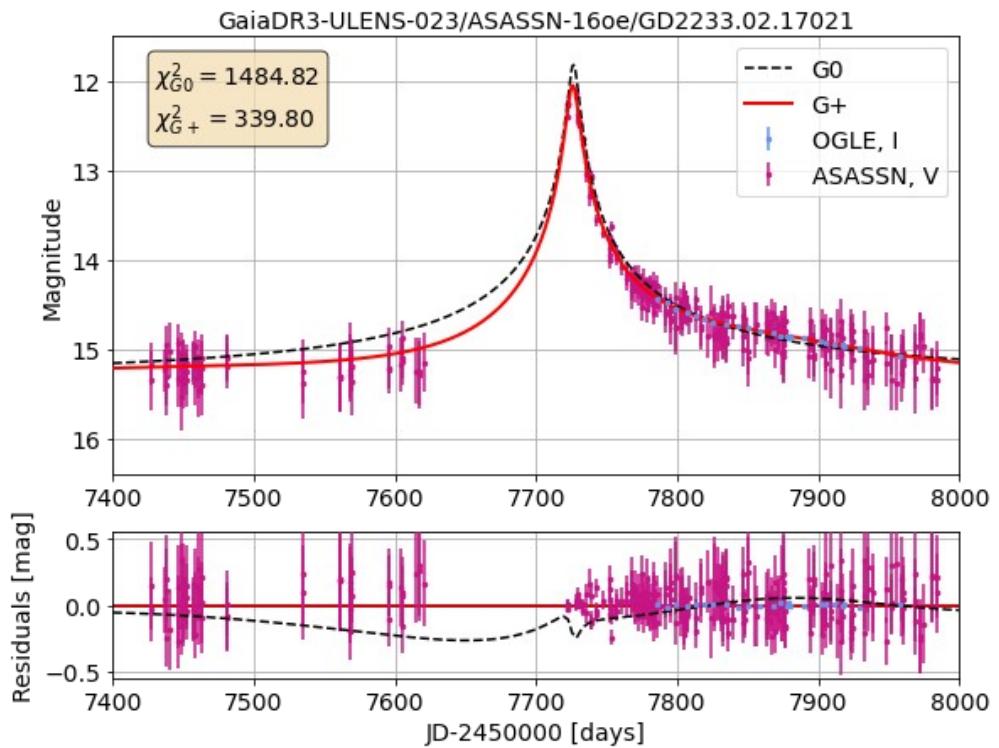
July 2014 – May 2017

Wyrzykowski, Kruszyńska et al.
submitted



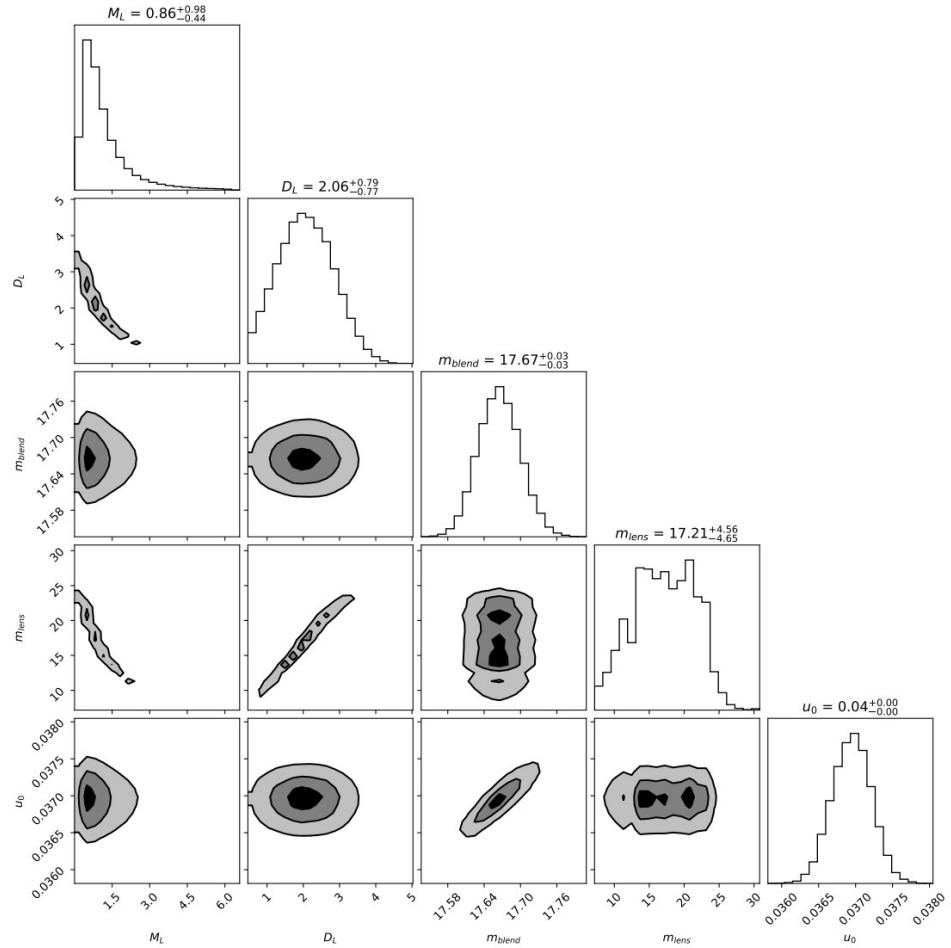
The future

- Gaia DR3 microlensing catalogue
- Applying the same method of mass and distance determination to all events from Gaia DR3...



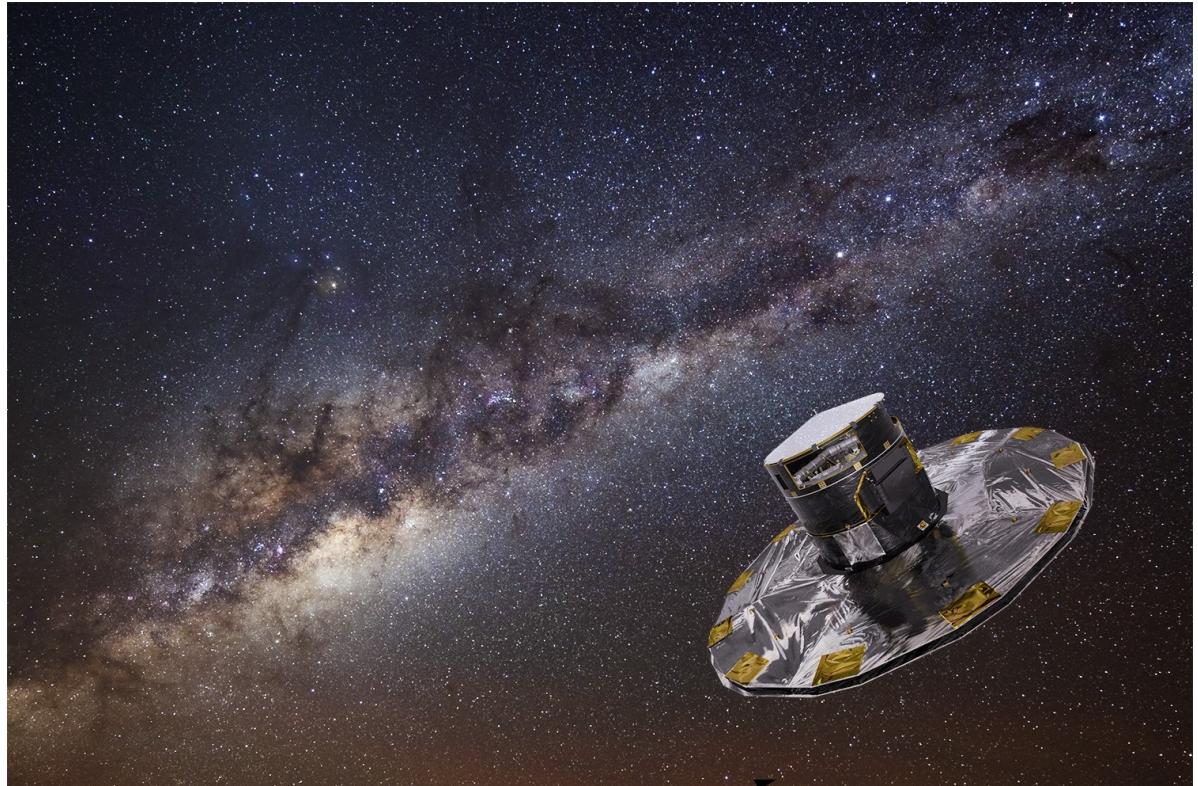
The future

- Gaia DR3 microlensing catalogue
- Applying the same method of mass and distance determination to all events from Gaia DR3...



The future

- Gaia DR3 microlensing catalogue
- Applying the same method of mass and distance determination to all events from Gaia DR3...
- And to events found in GSA
- Stay tuned!
- Also enjoy DR3~



Source: ESA/DPAC